



## **Reproducing the Profile of the Coriolis number of the Sun** Kayleigh Johnson 7 August 2020

## Introduction

- Layers of the Sun •
  - Core •
  - Radiative zone •
  - Convection zone
- Sun's radius-700 Mm ۲
  - Convection zone- 200 Mm •
  - NSSL-35Mm



 $10^{-3} \le Co(NSSL) \le 1 \le Co(CZ) \le 10$ 

 $Co = 2\tau \Omega$ 

6.0

r [10<sup>8</sup> m]



7.0





## Pencil Code

Table 1: Simulation parameters and their units set in the code.  $n_i$  is the number of mesh points.

Radius= 1.00 Width=0.02 Forcing=0.10

| parameter   | value   |
|-------------|---|
| $n_r$       | $128 \ (0.7 \le r \le 1)$                       |
| $n_{	heta}$ | $256 \ (72^{\circ} \le \theta \le 108^{\circ})$ |
| $n_{\phi}$  | $64~(0^{\circ} \le \phi \le 15^{\circ})$        |
| ν           | $10^{-5}$                                       |
| $c_s$       | 1   |
| $f_0$       | 0.02  |
| g           | 3   |



**Results** 



- With my simulation setup, I was able reproduce the radial dependence of the Coriolis number.
- The trend is correct, but the values are larger than what was expected.

 $10^{-3} \leq Co(NSSL) \leq 1 \leq Co(CZ) \leq 10$ 



- The radial dependency of the Coriolis number can be achieved with this simple set up
- For future work, I plan to continue with parameter studies allow for individual variables to be differed to obtain the expected Coriolis number profile
  - Parameters that I plan to study:
    - Forcing Width
    - Forcing Strength
    - Forcing Location
    - Rotation Rate



- I would like to thank Maarit Käpylä and Atefeh Barekat for helping me.
- I would also like to thank Dr. Erik Gilson for allowing me to work with him
- This work was made possible by funding from the Department of Energy for the Summer Undergraduate Laboratory Internship (SULI) program. This work is supported by the US DOE Contract No. DE-AC02-09CH11466.